My Question : please variation this action with respect to tensor metric  $g_{\mu\nu}$ .this called the Einstein equation. To obtain the Einstein equation, we vary the action with respect to the metric tensor

$$\begin{split} s &= \int d^4x \, \sqrt{-g} [R + \, \alpha(\varphi) R_{GB} \\ R_{GB} &= R^2 - 4 R_{\mu\nu} R^{\mu\nu} + R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} \end{split}$$

where R is the Ricci scalar, R\_{ $\mu\nu$ } is a Ricci tensor, R\_{ $\mu\nu\rho\sigma$ } is a Riemann tensor, g is the determinant of the metric tensor,  $\alpha(\phi)$  is a function of  $\phi(r)$ ,  $R_{GB}$  is a quadratic Gauss-bonnet (GB) term, G is the gravitational constant, and (d^4x) represents the volume element in four-dimensional spacetime.

$$\alpha(\phi) = f(\phi)$$
  
 $\phi = \phi(r)$ 

Let us now turn our attention to the third term  $\delta S_3$  in (19.34), in which we must express  $\delta \sqrt{-g}$  in terms of the variation  $\delta g^{\mu\nu}$ . Recalling that  $g=\det[g_{\mu\nu}]$ , we note that the cofactor of the element  $g_{\mu\nu}$  in this determinant is  $gg_{\mu\nu}$ . It follows that

$$\delta g = g g^{\mu\nu} \delta g_{\mu\nu} = -g g_{\mu\nu} \delta g^{\mu\nu},$$

where in the second equality we have used the result (19.33). Thus, we have

$$\delta \sqrt{-g} = -\frac{1}{2}(-g)^{-1/2}\delta g = -\frac{1}{2}\sqrt{-g}g_{\mu\nu}\delta g^{\mu\nu}.$$

I see different results in papers, but I don't have any code in Maple for them. I need the Maple code for this equation to continue deriving my equation.

so i give you some example of this result in next page:

Example 1: (arXiv:1812.06941v1 [hep-th] 17 Dec 2018)

$$S = \frac{1}{16\pi} \int d^4x \sqrt{-g} \left[ R - \frac{1}{2} \partial_{\mu}\phi \partial^{\mu}\phi + f(\phi) R_{GB}^2 - 2\Lambda \right].$$

$$R_{GB}^2 = R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma} - 4R_{\mu\nu}R^{\mu\nu} + R^2 \,, \label{eq:RGB}$$

by varying the action (S) with respect to the metric tensor  $g_{\mu\nu}$  , we derive the gravitationalfiel equations:

$$G_{\mu\nu} = T_{\mu\nu}$$

where  $G_{\mu\nu}$  is the Einstein tensor and  $T_{\mu\nu}$  is the energy-momentum tensor, with the latter having the form

$$T_{\mu\nu} = -\frac{1}{4} g_{\mu\nu} \partial_{\rho} \phi \partial^{\rho} \phi + \frac{1}{2} \partial_{\mu} \phi \partial_{\nu} \phi - \frac{1}{2} \left( g_{\rho\mu} g_{\lambda\nu} + g_{\lambda\mu} g_{\rho\nu} \right) \eta^{\kappa\lambda\alpha\beta} \tilde{R}^{\rho\gamma}_{\alpha\beta} \nabla_{\gamma} \partial_{\kappa} f(\phi) - \Lambda g_{\mu\nu} .$$

example2: (arXiv:hep-th/0504052v2 25 May 2005)

The starting action is given by

$$S = \int d^4x \sqrt{-g} \left\{ \frac{1}{2\kappa^2} R - \frac{\gamma}{2} \partial_\mu \phi \partial^\mu \phi - V(\phi) + f(\phi) G \right\} .$$

On the other hand, the variation over the metric  $g_{\mu\nu}$  gives

$$0 = \frac{1}{\kappa^2} \left( -R^{\mu\nu} + \frac{1}{2} g^{\mu\nu} R \right) + \gamma \left( \frac{1}{2} \partial^{\mu} \phi \partial^{\nu} \phi - \frac{1}{4} g^{\mu\nu} \partial_{\rho} \phi \partial^{\rho} \phi \right) + \frac{1}{2} g^{\mu\nu} \left( -V(\phi) + f(\phi) G \right)$$

$$-2f(\phi) R R^{\mu\nu} + 2\nabla^{\mu} \nabla^{\nu} \left( f(\phi) R \right) - 2g^{\mu\nu} \nabla^2 \left( f(\phi) R \right)$$

$$+8f(\phi) R^{\mu}_{\ \rho} R^{\nu\rho} - 4\nabla_{\rho} \nabla^{\mu} \left( f(\phi) R^{\nu\rho} \right) - 4\nabla_{\rho} \nabla^{\nu} \left( f(\phi) R^{\mu\rho} \right)$$

$$+4\nabla^2 \left( f(\phi) R^{\mu\nu} \right) + 4g^{\mu\nu} \nabla_{\rho} \nabla_{\sigma} \left( f(\phi) R^{\rho\sigma} \right) - 2f(\phi) R^{\mu\rho\sigma\tau} R^{\nu}_{\ \rho\sigma\tau} + 4\nabla_{\rho} \nabla_{\sigma} \left( f(\phi) R^{\mu\rho\sigma\nu} \right)$$